

BE-BI Options for Satellite, Ghost, Debunched Beam and Bunch Length Measurements – Part I/II

Ralph J. Steinhagen

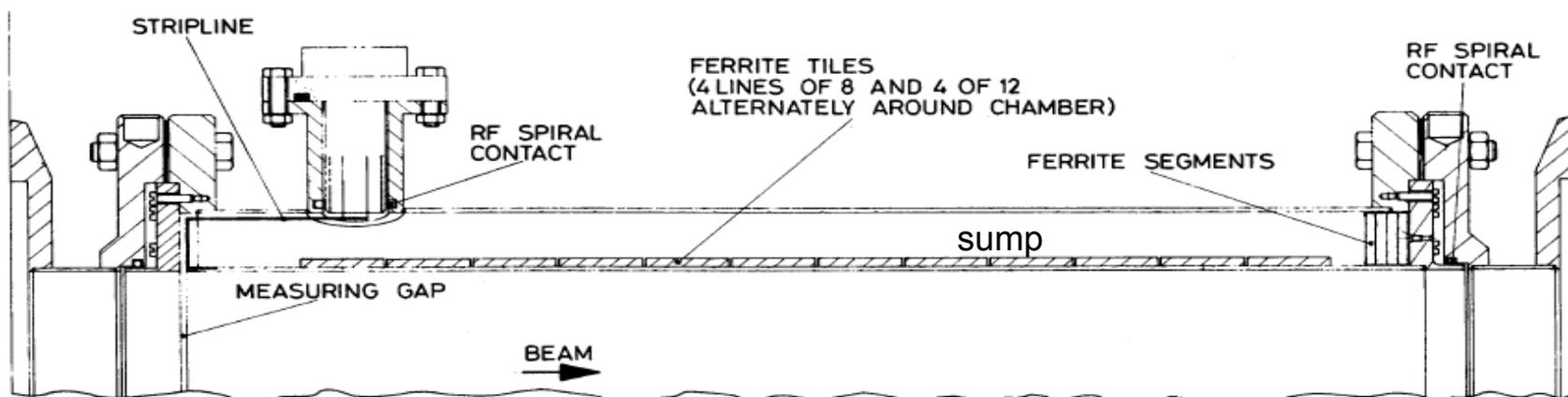
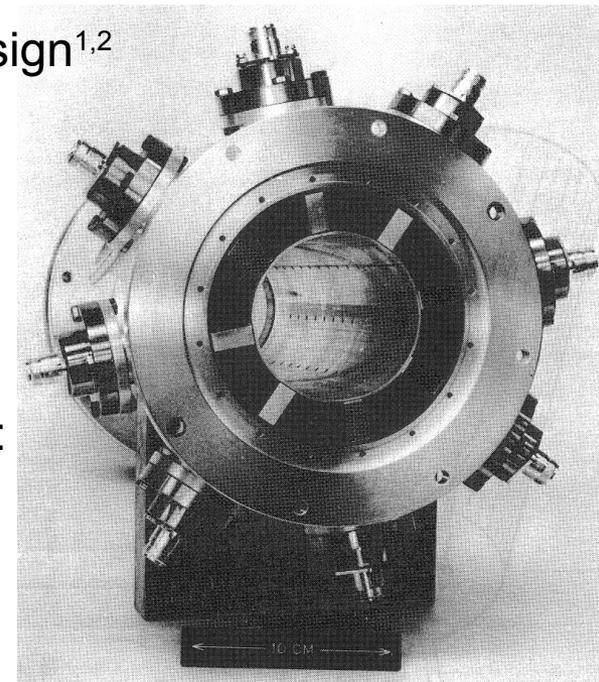
Special thanks to Th. Bohl, S. Bart-Pedersen and H. Damerau

Resumé:

- Detection of ~1%-level Satellites:
 - existing PS/SPS pick-up hardware (WCM) fulfills most requirements (except de-bunched beam detection)
 - 'Visually' easy to detect but ... fully automated 'turn-key' system requires system response compensation, further control room level integration
 - leverage experiences with LHC BI-WCM could be applied to PS/SPS
- Detection of sub-percent level Satellites ('ghosts') or un-bunched beam:
 - require/install new high-bandwidth, low-noise pick-ups
 - can re-use existing acquisition, post-processing and CCC integration

Wall Current Monitor as used by BI

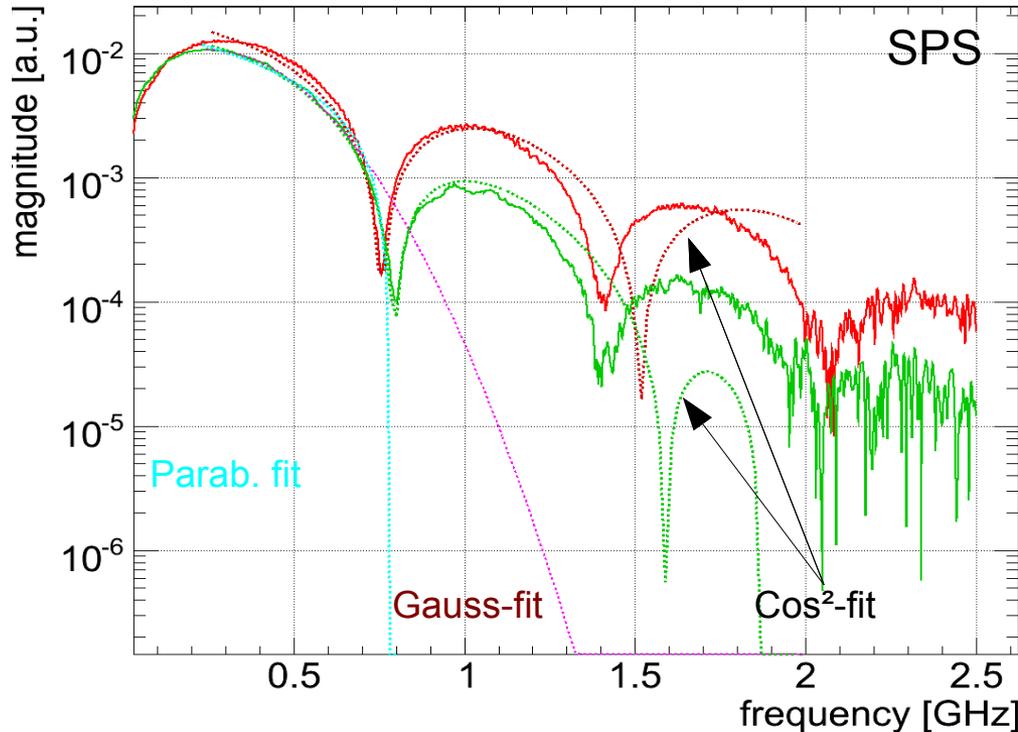
- LHC/SPS WCM pickup based on established 78' design^{1,2}
- Simplicity is key necessity to control systematics and reflections on the 10^{-3} level at GHz frequencies:
 - WCM + “combiner” → 3/8” → 30 (100) m 7/8” cable
 - 40 dB attenuator → 3 GHz fast sampling scope
 (N.B. Implies control of every single transition/bend/connector on mm-level)
- Idea was not to re-build the turn-based BQM system:
 - a) Tackling average signal over N-turns
 - overcomes scope quantisation/noise
 - b) full compensation of measured system response
 - necessary to get (any hope of) %-accuracy



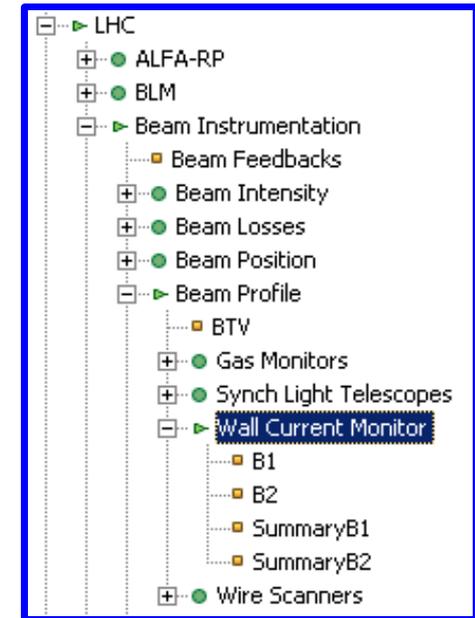
¹T. Linnear, “The high frequency longitudinal and transverse pick-ups used in the SPS”, CERN-SPS/ARF/78-17, 1978

²Th. Bohl, “The APWL Wideband Wall Current Monitor”, CERN-BE-2009-999, 2009

- Real bunches do not necessarily obey 'Gaussian' shapes

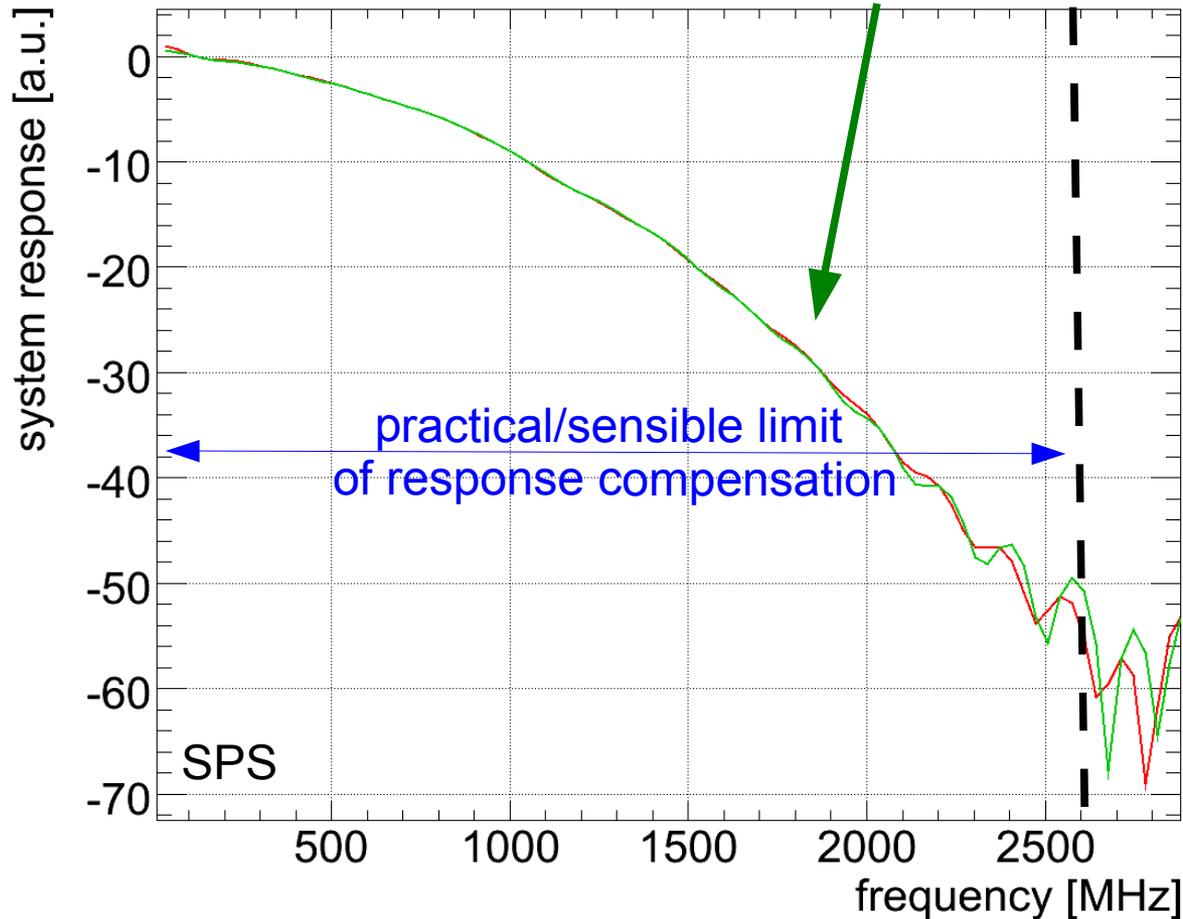


In Timber:



- What's being computed so far:
 - number & intensities of bunches & satellites (per 400 MHz bucket above thres.)
 - true Cos^2 - , Parabolic- & Gaussian bunch length χ^2 -fits
 - Frequency containing 50/95/99% of bunch power/intensities, peak voltages, ...
- Most difference/details are only visible at very high frequencies > 1 GHz
- Response of pick-up, cables, scope at these frequency need compensation!

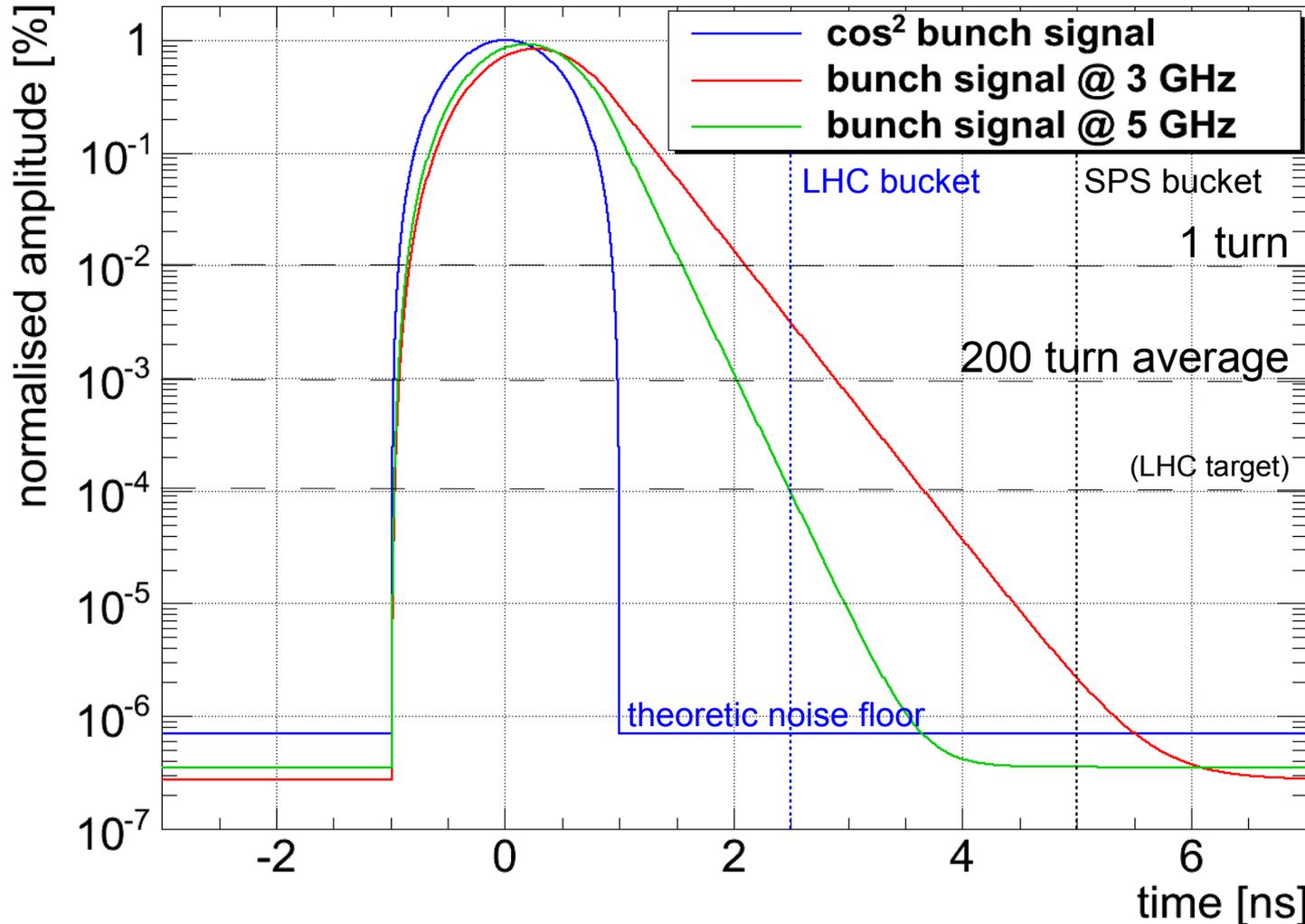
- True longitudinal bunch profile measurement is distorted by:
 - WCM pick-up response → design values + measurements by T. Bohl & U. Wehrle
 - combiner-response (star-topology) → only design (re-measure end '10)
 - Dispersion due to 7/8" Heliax cabling & analogue scope bandwidth



- Historical: (very) high numerical complexity if treating raw 20 (100) us frames

Fundamental limits of the WCM-based Scheme: 'Satellite' → 'Ghost' Detection Potential

- ... limited by total system bandwidth for below percent-level detection:

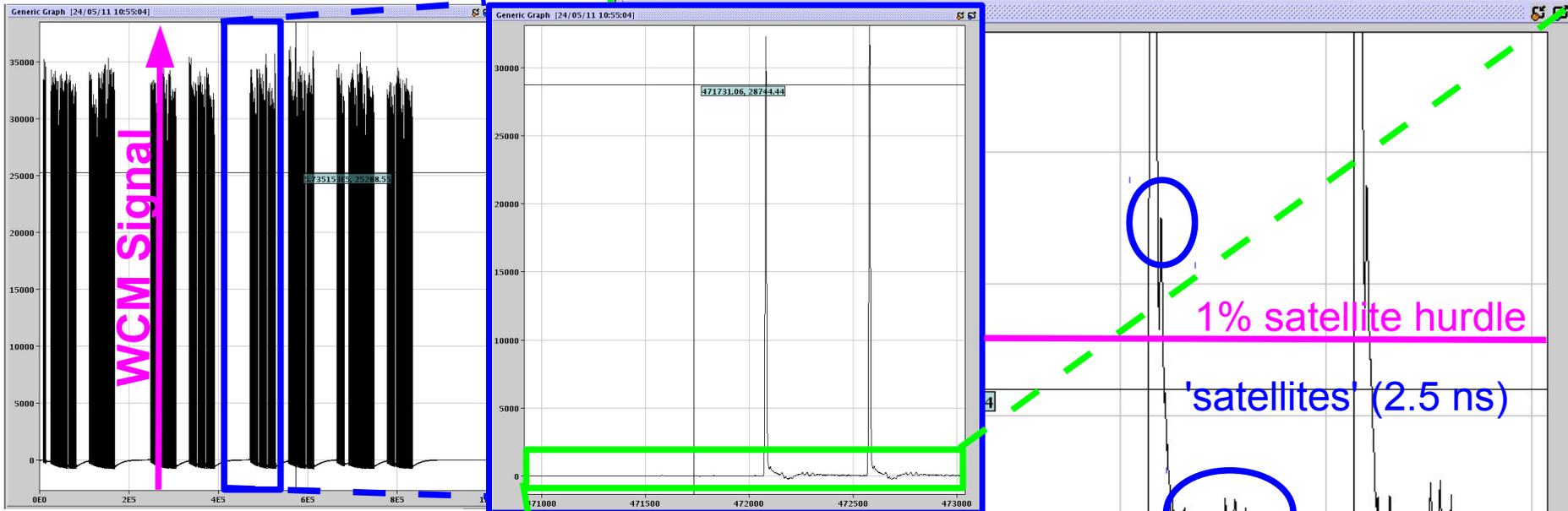


- ... limited by unavoidable systematic due to transmission line transitions, reflections, etc. (N.B. difficult to control better than 10^{-3} on > 2 m distances)

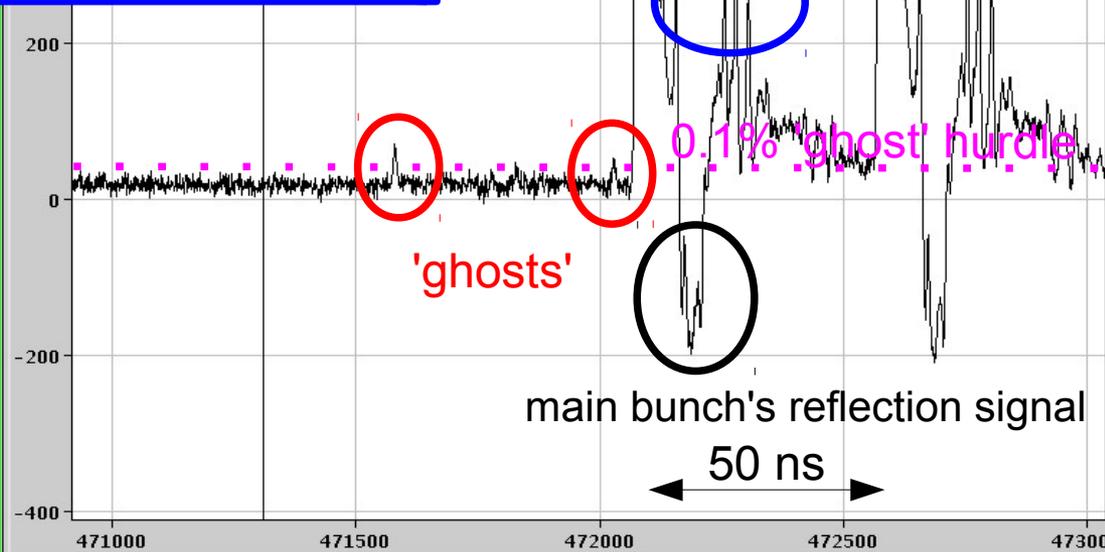


LHC Wall-Current-Monitor (based on old SPS RF design)

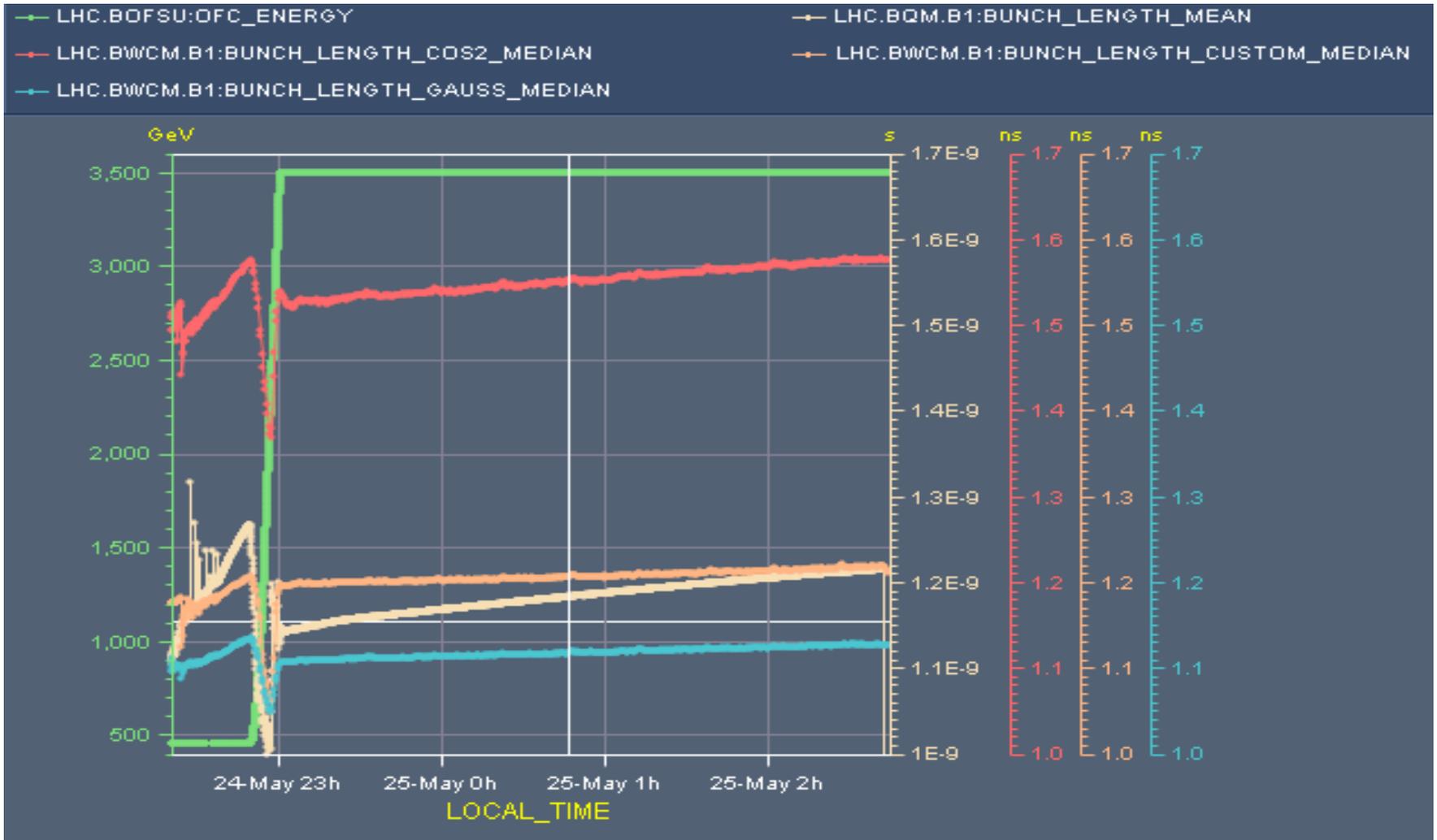
- LHC setup: WCM → short (30 m) 7/8" cable → 3 GHz Scope → post-processing...



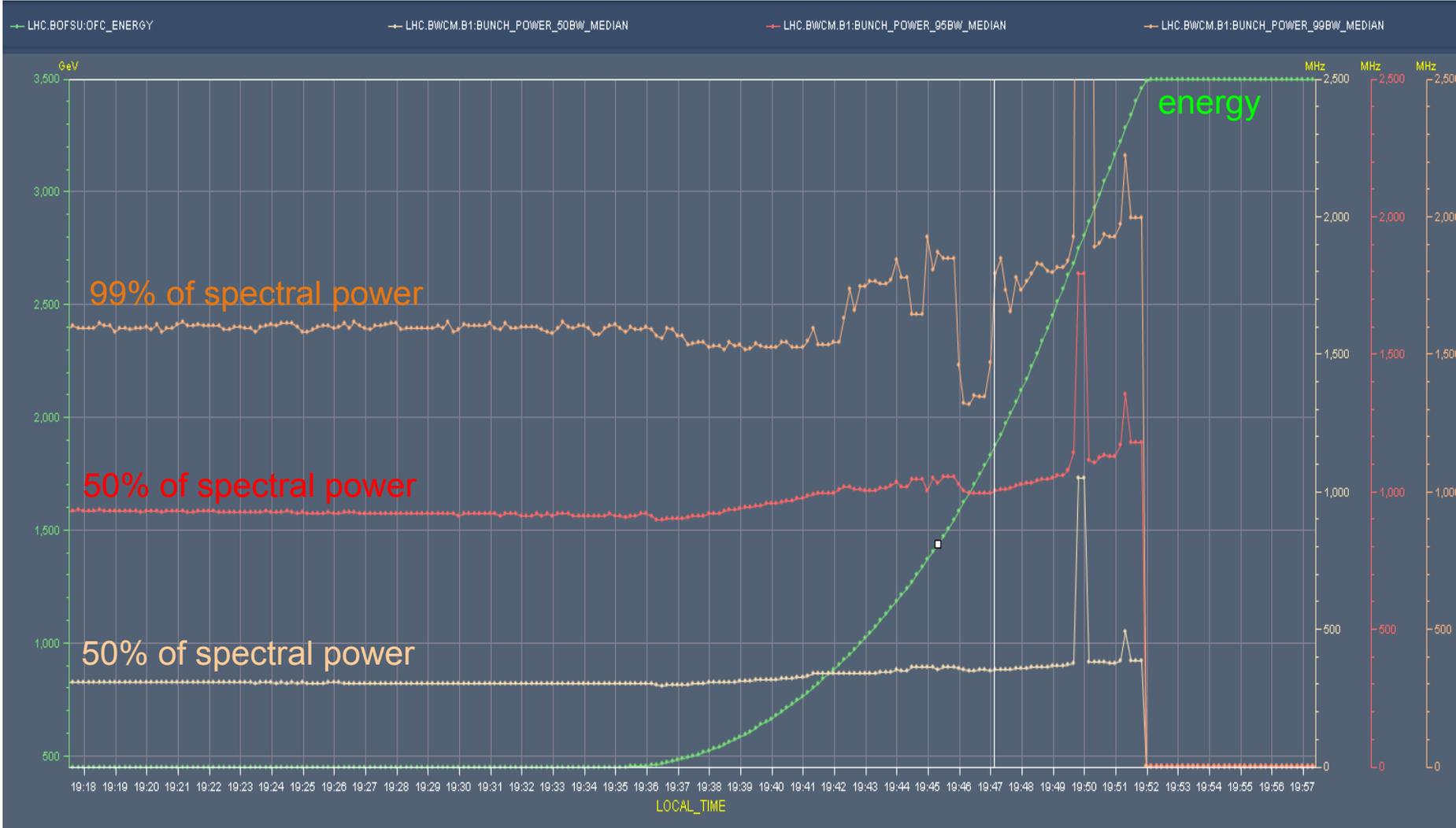
- Pick-up based on 70ies SPS RF design
- BI's mode of operation: 200 turn average + response compensation
- Limited by systematics (reflections, tails, etc.)



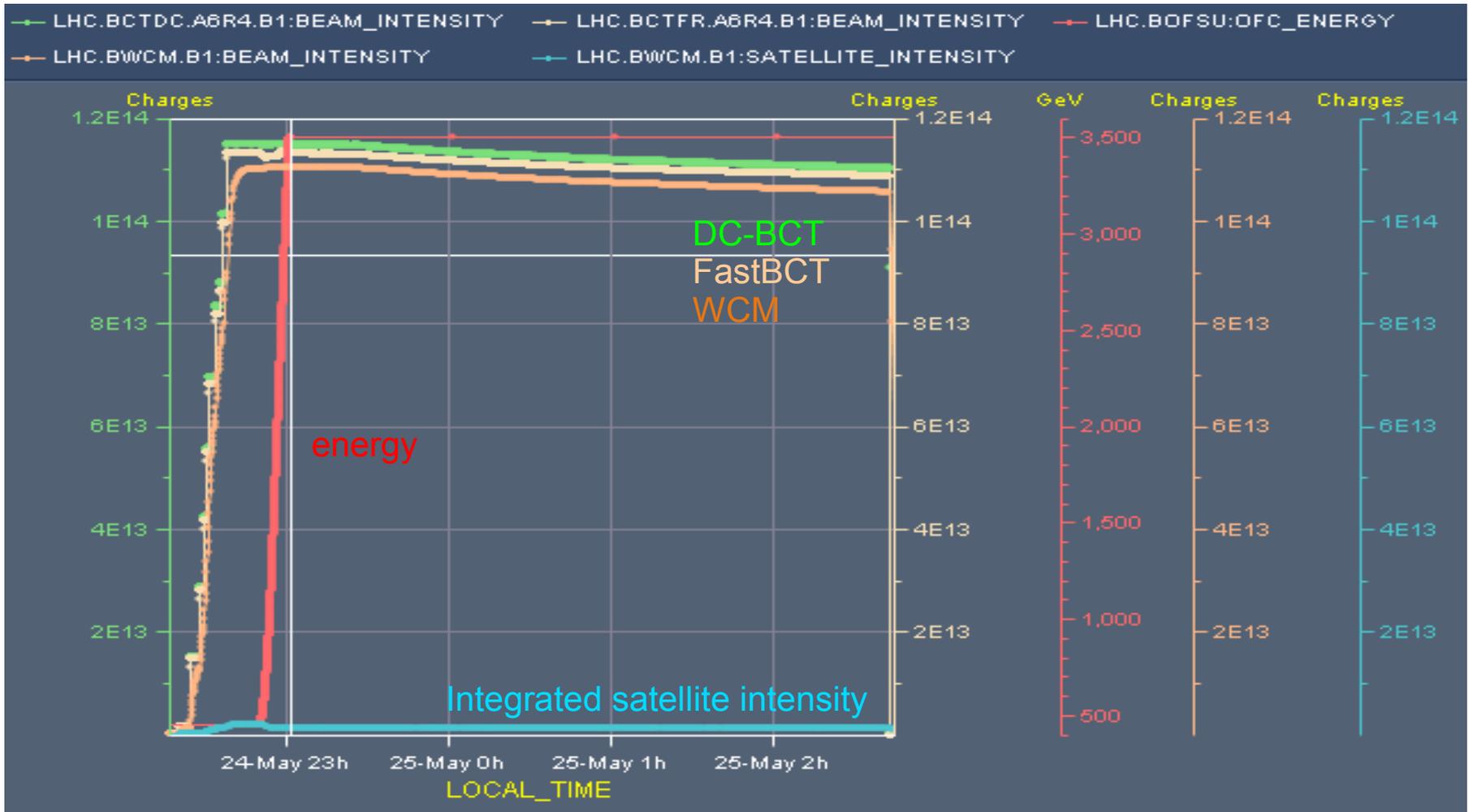
Comparison of Bunch Length Estimates



- ... there is no obvious bunch length → shape changes are important
 - difference between FWHM (BQM) and x^2 -fit Gaussian length estimate

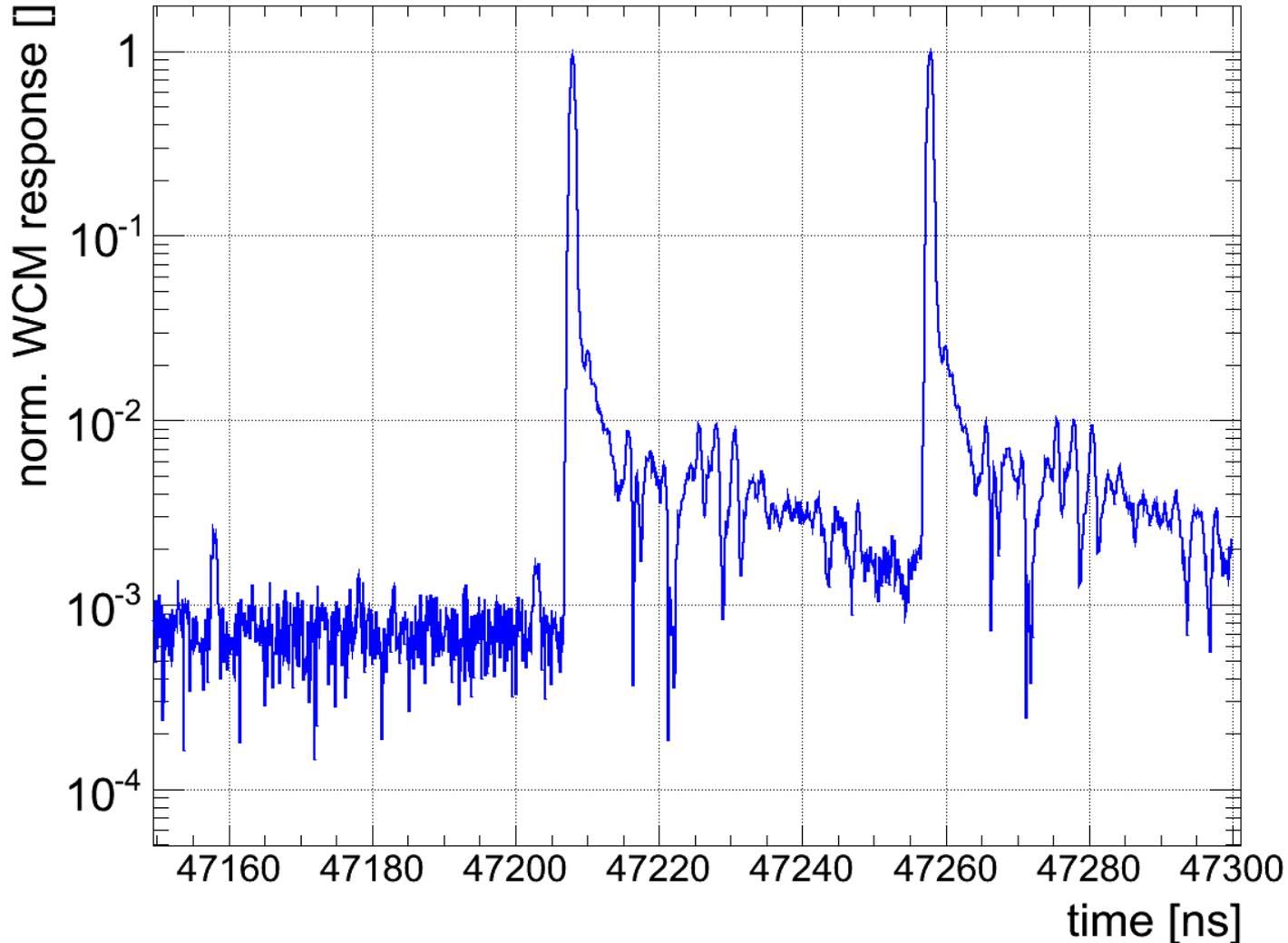


- Estimates give an indication of shape and required device bandwidths



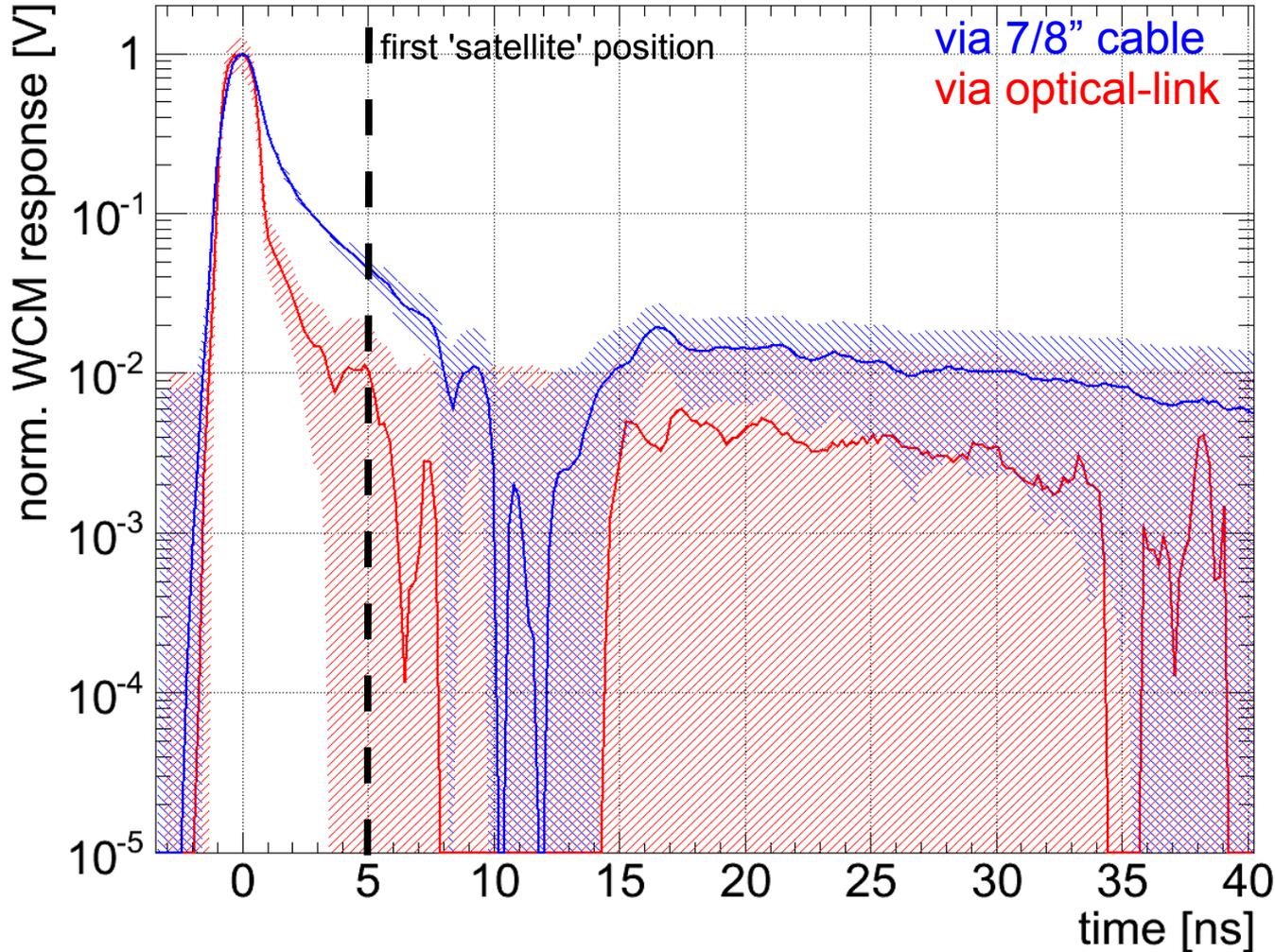
- WCM cross-calibrated to DC-BCT using a single nominal bunch (satellite free)
 - Typically percent-level beam outside nominal bucket
- Being addressed: local 400 MHz phase stability → affects 1st satellite after main bunch

- Example: satellites 50 (PS?) and 2.5 ns (LHC) prior to bunch train



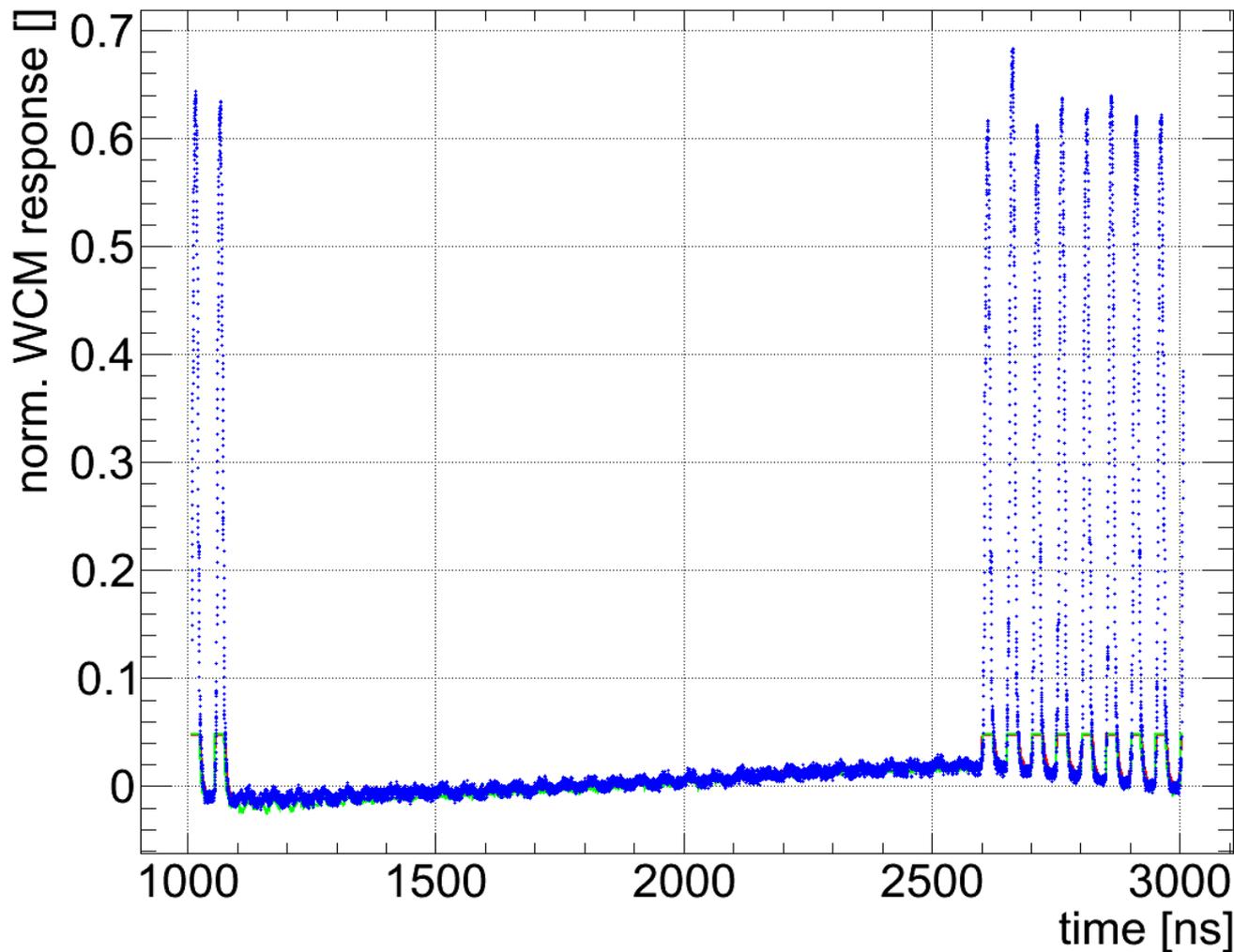
- 2.5 ns satellites after bunch visible but dominated by WCM tails/reflections...

- “Mother” design for LHC APWL, would expect similar performance



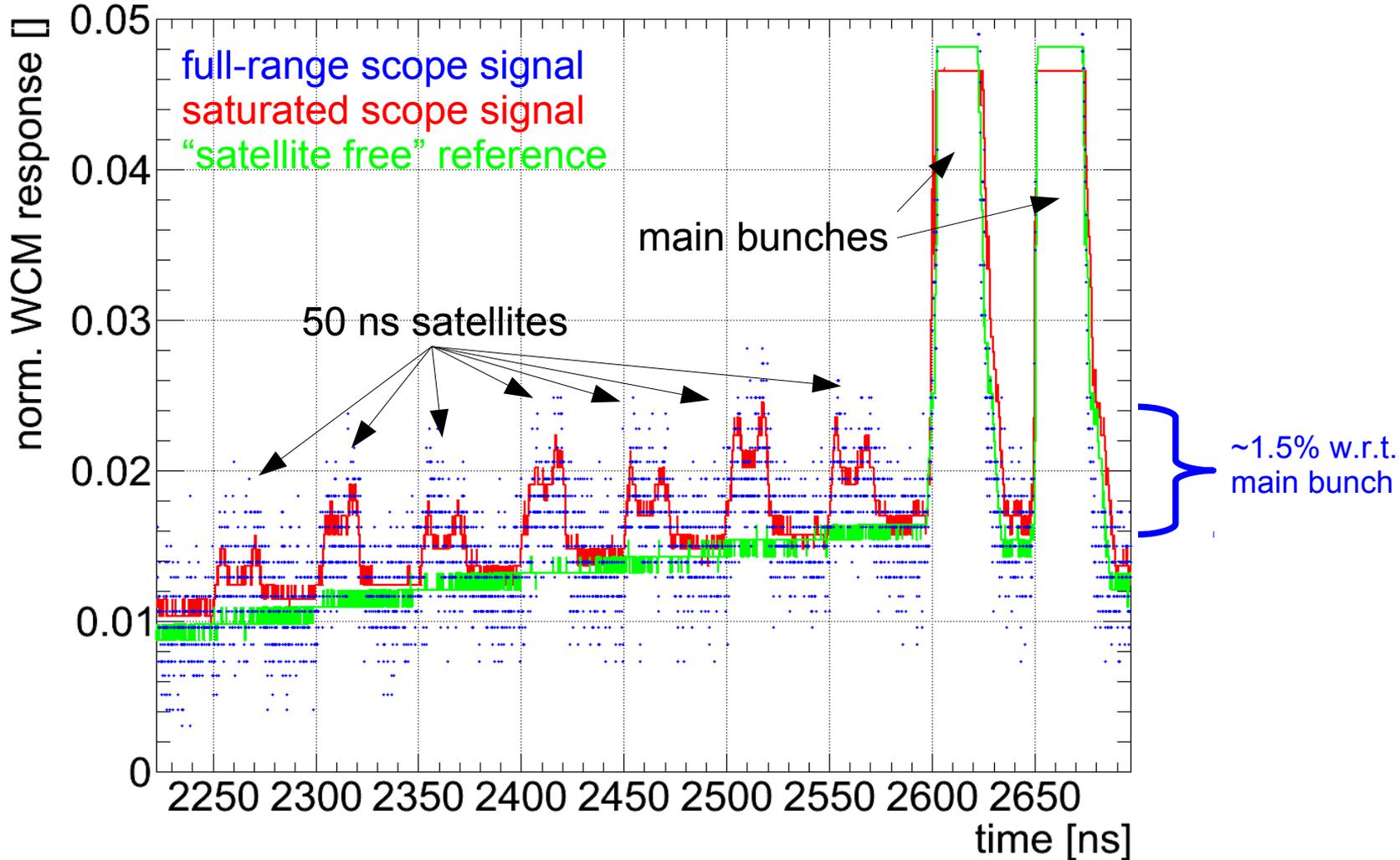
- higher bandwidth with optical link but noise compared to 7/8" cable
 → shorter cables/acquisition system in the SPS tunnel needed

- Initial test comparing single turn acquisition (no 200 turn avg. yet ... being analysed)



- Dominated by WCM systematic, known tails & reflections → upgrade planned

- Forcing satellites and saturating the scope input (fast recovery time)



- Satellites 'visible' and results look promising but requires post treatment to compensate for reflections, pick-ups response, droop etc.

Following slides focus more on the technical implementation aspect

- Finite Estimates (fit-limits <2.5 ns around peak or 3x noise-floor)
 - COS²-Distribution (probably best): $f(t) = I \cdot \frac{2}{B} \left[\cos\left(\pi \frac{t}{B}\right) \right]^2$ for $t \in [-B/2, +B/2]$, 0 elsewhere
 - BUNCH_LENGTH_COS2 (DB)/
bunchLengthCOS2 (FESA?)
 - Parabolic-distribution:
 - BUNCH_LENGTH_PARABOLIC/
bunchLengthParabolic
 - 50/95/99% power (by-product of deconvolution/intensity estimate)
 - BUNCH_LENGTH_POWER50, BUNCH_LENGTH_POWER90.../
bunchLengthPower50, bunchLengthPower95,
- Infinite estimates (N.B. non-physical since RF bucket is finite < 2.5 ns)
 - Full-Width-Half-Maximum (see plot): $FWHM = |x_2 - x_1|$
 - Gaussian distribution: $f(x) = \frac{1}{\sqrt{2\pi}\sigma_t} \cdot e^{-\frac{1}{2}\left(\frac{t-\mu}{\sigma_t}\right)^2}$
 - BUNCH_LENGTH_GAUSS/
bunchLengthGauss
 - RMS (alternate to Gaussian)
 - BUNCH_LENGTH_RMS/
bunchLengthRMS

$$\sigma = \langle x^2 \rangle - \langle x \rangle^2 \approx \frac{1}{N} \sum_i^N (x_i - \mu)^2$$

